

Fig. 33 shows the circuit of a simple test equipment that will allow you to measure signals that are beyond the normal bandwidth of an inexpensive oscilloscope. This circuit was built to monitor the frequency of a 10-m signal on a scope with a 5-MHz upper-frequency limit. The circuit features a Mini-Circuits Laboratory SRA-1 mixer. Any stable oscillator with an output of 10 dBm can be used for the local oscillator (LO), which is mixed with the HF signal to produce an IF within the bandwidth of the oscilloscope. The mixer can handle RF signal levels up to +3 dBm without clipping, so this was chosen as an upper limit for the RF input. A 1:1 transformer coupler is constructed by winding a 31-turn secondary of #28-gauge wire on an FT-37-75 core for RG-58 or an FT-50-75 core for use with larger coax such as RG-58. The primary is the piece of coaxial cable that goes through the core center. The coupler gives 30 dB of attenuation and has a frequency response from 0.5 to 100 MHz. An additional 20-dB of attenuation was added for a total of 50 dB before the mixer. One-percent resistors will do fine for the attenuator. The completed adapter should be built in a shielded box.

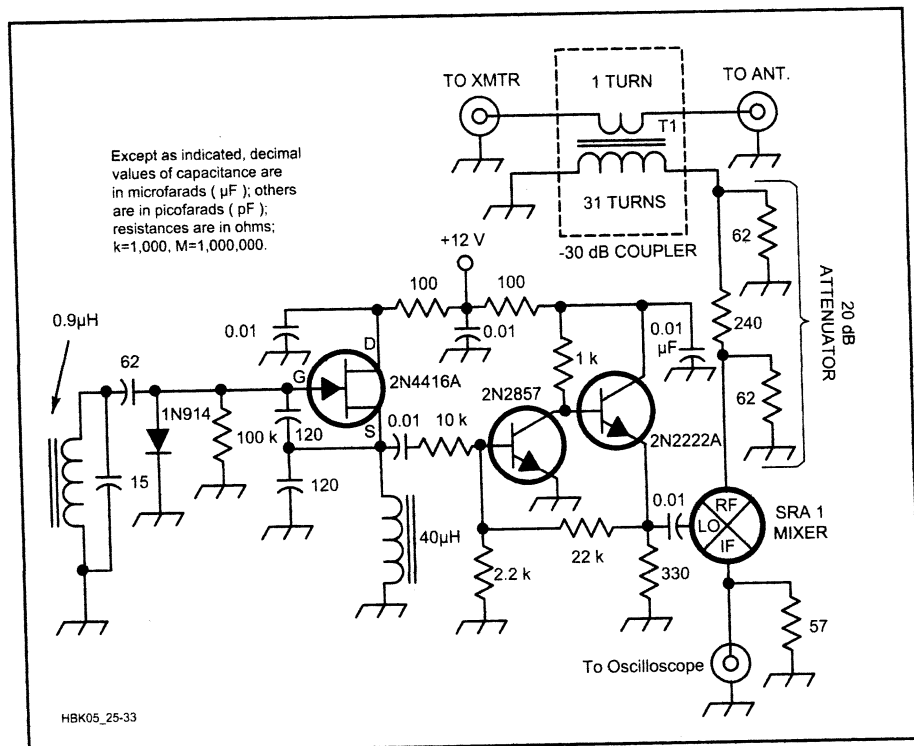


Fig 25.33 — This adapter displays HF signals on a narrow-bandwidth oscilloscope. It uses a 10-dBm 25-MHz LO, -30-dB coupler, 20-dB attenuator and diode-ring mixer. See text for further information.

quency displayed will be the difference between the LO and the input signal. As an example, a 28.1-MHz input and a 25-MHz LO will be seen as a 3.1-MHz signal on the oscilloscope.

More attenuation will be required for higher-power transmitters. This circuit was described by Kenneth Stringham Jr, AE1X, in the Hints and Kinks column of February 1982 *QST*.

A CALIBRATED NOISE SOURCE

NOISE FIGURE MEASUREMENT

One of the most important measurements in communications is the noise figure of a receiving setup. Relative measurements are often easy, while accurate ones are more difficult and expensive. One EME (moon bounce) station checks noise and system performance by measuring the noise of the sun reflected off the moon. While the measurement source (use of the sun and moon) is not expensive, the measuring equipment on 2 m consists of 48 antennas (each over 30 ft long). This measurement equipment is not for everyone!

The rest of us use more conventional noise sources and measuring techniques. Coverage of noise figure and its measurement appear in the **Receivers and Trans-**

mitters chapter of this *Handbook*.

Most calibrated and stable noise sources are expensive, but not this unit developed by Bill Sabin, WØIYH. It first appeared in May 1994 *QST*. When hams use a noise source, it is usually included in an RF bridge used to measure impedances and adjust antenna tuners. A somewhat different device (an *accurately calibrated and stable* noise source) is also useful. Combining a broadband RF noise source of known power output and a known output impedance with a true-RMS voltmeter, results in an excellent instrument for making interesting and revealing measurements on a variety of circuits hams commonly use. (Later on, some examples will be described.) The true-RMS voltmeter can be

an RF voltmeter, a spectrum analyzer or an AF voltmeter at the output of a linear receiver.¹

Calibrated noise generators and noise-figure meters are available at medium to astronomical prices. Here is a low-cost approach which can be used with reasonable confidence for many amateur applications where accuracy to tenths of a decibel is not needed, but where precision (repeatability) and comparative measurements are much more important. PC boards are available for this project.²

Semiconductor Noise Diodes

Any Zener diode can be used as a source of noise. If, however, the source is to be calibrated and used for reliable mea-